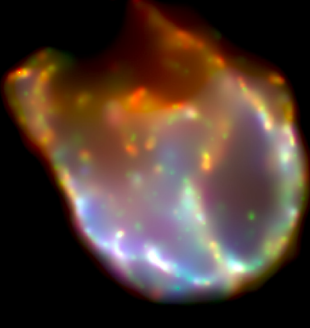
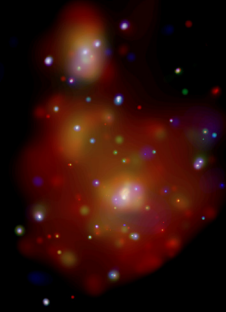
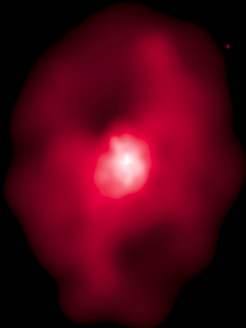
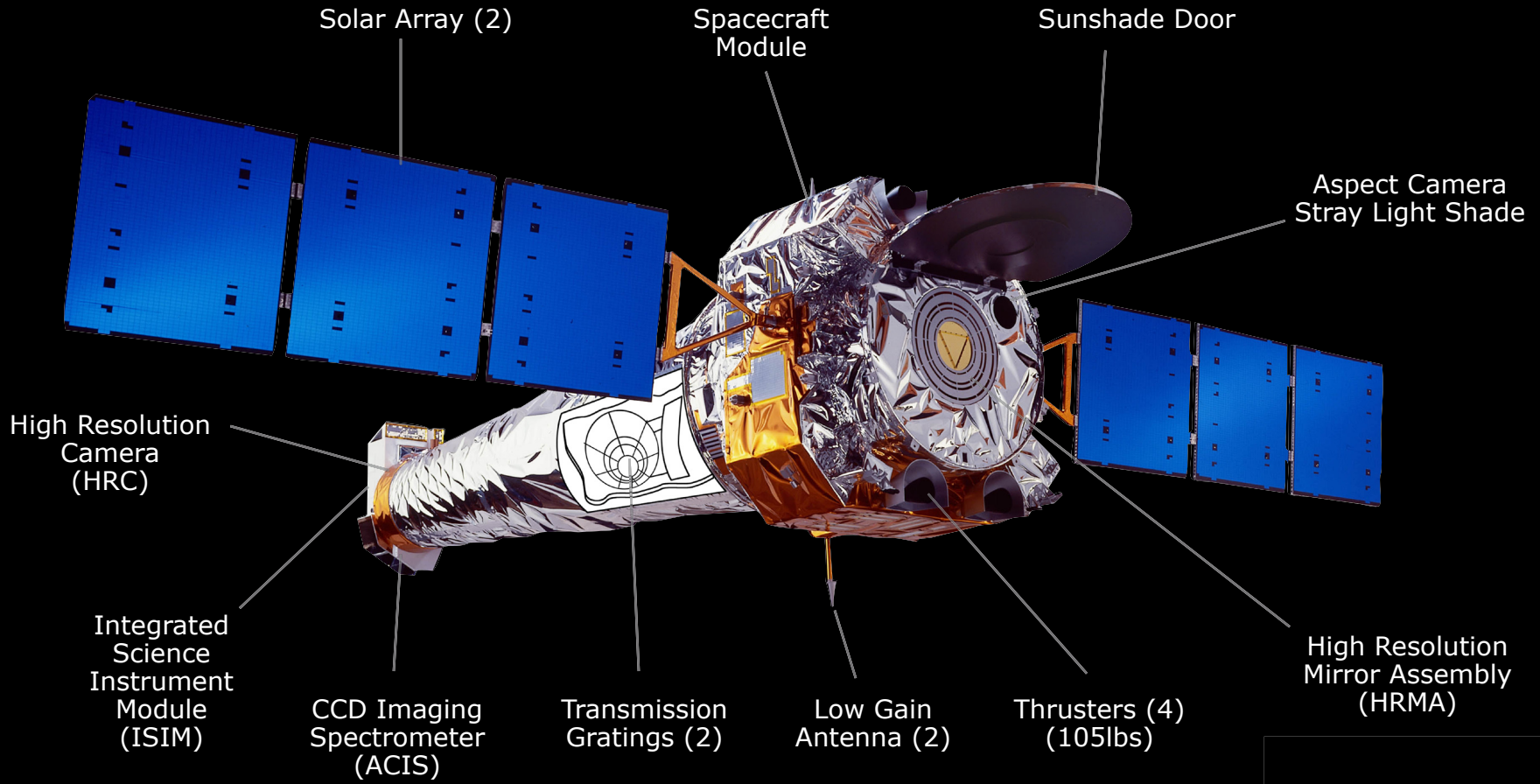
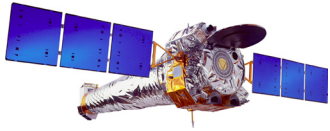


# CHANDRA SPECIFICATIONS



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*An X-ray telescope is the only way astronomers can observe the hot regions of the Universe. The most powerful optical telescopes, such as the Hubble Space Telescope, cannot see the vast clouds of hot gas that stretch millions of*

*light years across and contain enough matter to make hundreds of trillions of stars. X-ray telescopes allow us to image matter swirling as close as 90 kilometers from the event horizon of a stellar black hole or to track the expansion of a hot gas bubble produced by an exploding star.*

*Chandra is the third of NASA's Great Observatories. The mirrors on Chandra are the largest, most precisely shaped and aligned, and smoothest mirrors ever constructed. The images Chandra makes are twenty-five times sharper than the best previous X-ray telescope. Chandra, which was launched by the Space Shuttle on July 23, 1999, is helping scientists to better understand the hot, turbulent regions of space and answer fundamental questions about the origin, evolution, and destiny of the Universe.*

## Overall Specifications

Size (solar arrays deployed): 13.8 m x 19.5 m (45.3 ft x 64.0 ft)  
Weight: 4,800 kg (10,560 pounds)  
Orbit: 10,000 km x 140,161 km (6,200 x 86,900 miles);  
28.5 degree inclination  
Ascending node: 200 degrees  
Argument of perigee: 270 degrees  
Life: expected 10 years

## Spacecraft Specifications

Power: two 3-panel silicon solar arrays (2350 W)  
three 40 amp-hour nickel hydrogen batteries  
Antennas: two low-gain, conical log spiral antennas  
Frequencies: transmit 2250 MHz, receive 2071.8 MHz  
Command Link: 2 kilobits per second (kbps)  
Data Recording: solid state recorder; 1.8 gigabits  
(16.8 hours) recording capability  
Downlink Operations: downloaded typically every 8 hours  
Contingency Mode: 32 kbps  
Safing: autonomous operation

## Science Instruments

Advanced Charged Couple Imaging Spectrometer (ACIS): Ten CCD chips in 2 arrays provide imaging and spectroscopy; imaging resolution is 0.5 arc-sec over the energy range 0.2 - 10 keV; sensitivity:  $4 \times 10^{-15}$  ergs-cm<sup>-2</sup> sec<sup>-1</sup> in 10<sup>5</sup> s  
High Resolution Camera (HRC): Uses large field-of-view micro-channel plates to make X-ray images: ang. resolution < 0.5 arc-sec over field-of-view 31x31 arc-min; time resolution: 16 micro-sec. sensitivity:  $4 \times 10^{-15}$  ergs-cm<sup>-2</sup> sec<sup>-1</sup> in 10<sup>5</sup> s  
High Energy Transmission Grating (HETG): To be inserted into focused X-ray beam; provides spectral resolution of 60-1000 over the energy range 0.4 - 10 keV  
Low Energy Transmission Grating (LETG): To be inserted into focused X-ray beam; provides spectral resolution of 40-2000 over the energy range 0.09 - 3 keV

## Telescope System

High Resolution Mirror Assembly: 4 nested pairs of grazing incidence paraboloid and hyperboloid mirrors  
Length of Mirrors: each 83.3 cm (32.8 in) long  
Weight of Mirrors: 956.4 kg (2,104 pounds) total  
Focal Length: 10 meters (32.8 ft)  
Outer Diameter: 1.2 meters (3.9 ft)  
Field of View: 1.0 degree diameter  
Ang. Resolution: 0.5 arc-sec  
Altitude Control: 6 reaction wheel control  
2 inertial reference units  
Aspect Camera: 1.40 deg x 1.40 deg field-of-view  
Pointing Stability: 0.25 arc-sec (RMS) radius over 95% of all 10 second periods  
Pointing Accuracy: 30 arc-sec 99% of viewing time  
Remarks: Mirrors have an effective area of 400 sq. cm.  
@1 keV; 600 A iridium coating

The Chandra program is managed by the Marshall Spaceflight Center for NASA's Science Mission Directorate. Northrop Grumman (formerly TRW), the prime contractor, assembled and tested the observatory for NASA. The Chandra X-ray Center is operated for NASA by the Smithsonian Astrophysical Observatory.